

Signal to Noise Ratio(SNR) improvement of atmospheric signals using Variable windows

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Abstract:

Windows reduce the sidelobe leakages from the spectrum when applied to finite length time domain sequence. Adjustable windows can be used to control the amplitude of sidelobes with respect to main lobe. This paper presents the effect of window parameter in Dolph-chebyshev, Kaiser, Gaussian and Tukey windows on the Signal to Noise Ratio(SNR) of radar returns and proposes an optimum value of window parameters with which data may be weighed. A comparative study is made on the improvement of SNR using these variable windows.

Keywords:

Dolph-chebyshev, Kaiser, Gaussian and Tukey windows, sidelobe leakages

1. INTRODUCTION

Generally windows are time-domain weighting functions used for reduction Gibbs' oscillations resulting from the truncation of a Fourier series. The conventional method that is adopted is to multiply the data with a rectangular window of the same length as that of the data, finding Discrete Fourier Transform of the multiplied sequence and from this finding the power spectrum. The application of FFT to a finite length data leads to spectral leakage, picket fence effects and spectrum broadening. Therefore a window which gives a spectrum that has minimum leakage is naturally preferred. These effects can be reduced by weighting the data with suitable windows. Therefore, the problem associated with the spectral estimation of a finite length data by the FFT techniques is choosing efficient windows or data smoothing schemes for particular type of signals. Windows are used to weigh complex time series of the in-phase and quadrature components of the radar return samples prior to applying the DFT. The observed Doppler spectra therefore represent convolution of the Fourier transforms of the original signals with those of the weighting windows, projected onto the discrete (angular) frequencies.

Lot of work is reported in the literature on window functions. [1,2,3,4]. Windows can be categorized as fixed type windows and adjustable type windows[5,6]. Fixed type windows have one independent parameter, that is, the window length which controls the main-lobe width. Adjustable windows have two or more independent parameters, that is, window length, as in

fixed windows, and one more additional parameters that can control window characteristics [7]. Kaiser window has two parameters and achieves close approximation to discrete prolate functions that has maximum energy concentration in the main lobe [7, 8]. The Dolph-Chebyshev window has also two parameters and produces minimum main-lobe width for a specified maximum side-lobe level [7, 9,10]. The Kaiser and Dolph-Chebyshev windows are able to control the amplitude of the side lobes relative to that of the main lobe level. Through proper choice of these parameters, the amplitude of the side lobes relative to that of the main lobe can be controlled easily[12,13].

The radar returns is considered to be composed of monotonic atmospheric signal superimposed with white noise [15,16]. Since the signal does not correspond exactly to the sampling frequencies, the signal portions of the spectra follow those of the envelopes of the side lobe maxima [14]. Spectral leakage from the signal therefore exceeds noise level, evaluated by the method of Hildebrand and Sekhon, and a corresponding underestimate of signal-to-noise ratio (SNR). Several others tried to improve SNR using other methods such as parametric and group delay methods [21,22]

2. WINDOWS IN SPECTRUM ANALYSIS

Any signal can be described in different systems and there is always an engineering value in examining the signal on a basis system. One such basis system is the set of complex exponentials and the attraction of this basis set exponentials is that complex exponentials are the eigen-functions and eigen-series of linear time invariant differential and difference operators respectively. It means that when a sine wave is applied to an linear filter the steady state system response is a scaled version of the same sine wave. The system can affect the complex amplitude i.e. magnitude and phase of the sine wave but it can't change the frequency. Consequently complex sinusoids have become a standard tool to probe and describe linear time invariant (LTI) systems. The Fourier Transform and Fourier series, shown on the left and right hand side of (1), permits us to describe signals equally well in both the time domain and the frequency domain [4].

$$\left. \begin{aligned}
 H(\omega) &= \int_{-\infty}^{+\infty} h(t)e^{-j\omega t} dt \\
 h(t) &= \frac{1}{2\pi} \int_{-\infty}^{+\infty} H(\omega)e^{+j\omega t} d\omega \\
 H(\theta) &= \sum_{-\infty}^{+\infty} h(n)e^{-j\theta n} \\
 h(n) &= \frac{1}{2\pi} \int_{-\pi}^{+\pi} H(\theta)e^{+j\theta n} d\theta
 \end{aligned} \right\} \dots\dots\dots (1)$$

Since the complex exponentials have infinite support, the limits of integration in the forward transform (time- to- frequency) are from minus to plus infinity. As observed earlier, all signals of engineering interest have finite support; hence the limits of integration and summation of the

Fourier transform may take finite form by defining finite support window to limit the signal to the appropriate finite support interval as given in (2)

$$\left. \begin{aligned} H_w(\omega) &= \int_{-\infty}^{+\infty} w(t) \cdot h(t) e^{-j\omega t} dt \\ h(t) &= \frac{1}{2\pi} \int_{-\infty}^{+\infty} H_w(\omega) e^{+j\omega t} d\omega \\ H_w(\theta) &= \sum_{-\infty}^{+\infty} w(n) \cdot h(n) e^{-j\theta n} \\ h(n) &= \frac{1}{2\pi} \int_{-\pi}^{+\pi} H_w(\theta) e^{-j\theta n} d\theta \end{aligned} \right\} \dots\dots\dots (2)$$

Here, w(t) and w(n) are window functions in CT and DT domains. When examining (2), it may arise that to what extent the transform of the signal is limited with the multiplicative window. It is well known that multiplication of any two functions in the time domain is equivalent to convolution of the same functions spectra in the frequency domain. As shown in (2.3), the transform of windowed signal is the convolution of the transform of the signal with that of the transform of the window [11,17].

$$\left. \begin{aligned} H_w(\omega) &= \frac{1}{2\pi} \int_{-\infty}^{+\infty} H(\lambda) \cdot W(-\lambda) d\lambda \\ H_w(\theta) &= \frac{1}{2\pi} \int_{-\pi}^{+\pi} H(\lambda) \cdot W(\theta - \lambda) d\lambda \\ H(\omega) &= \frac{1}{2\pi} \int_{-\infty}^{+\infty} h(t) e^{-j\omega t} d\omega \\ W(\omega) &= \frac{1}{2\pi} \int_{-T/2}^{+T/2} w(t) e^{-j\omega t} d\omega \\ H(\theta) &= \sum_{-\infty}^{+\infty} h(n) e^{-j\theta n} \\ W(\theta) &= \sum_{-N/2}^{+N/2} w(n) e^{-j\theta n} \end{aligned} \right\} \dots\dots\dots (3)$$

The side-lobe structure of the windowed transform limits the ability of the transform to detect spectral components of significantly lower amplitude in the presence of a large amplitude

component while the main-lobe width of the windowed transform limits the ability of the transform to resolve or separate nearby spectral components. The first of these limitations is demonstrated in Figure 1 where a stylized power spectrum of two sinusoids of infinite extent and of finite extent is presented. From the Figure 1 it can be observed that the side-lobe level of the high level signal is greater than main lobe level of the low-level signal, and hence high level side lobe masks the presence of the low-level signal. If the low-level signal is to be detected in the presence of nearby high level signal, window applied to the signal has to be modified[18,19,20].

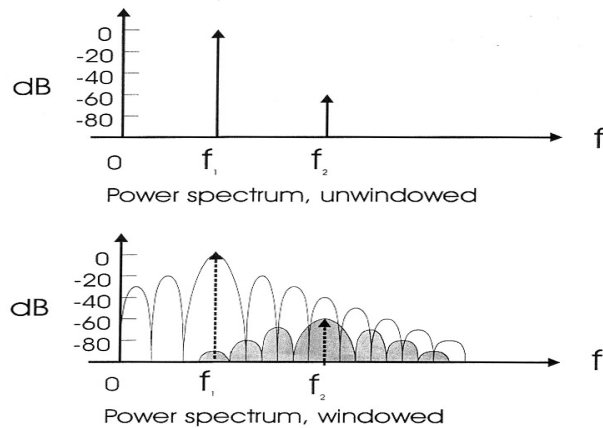


Figure:1 Spectral Representation of Un-windowed and of a Rectangle Windowed Sinusoids of Significantly Different Amplitudes.

3. VARIABLE WINDOWS APPLIED TO ATMOSPHERIC SIGNALS

In contrast to the fixed windows, the variable windows have two parameters: the length of the sequence and a shape parameter. In case of MST Radar the length of the window is fixed to 512 data points, only the shape parameter is varied. As the shape parameter changes the side lobe level alters. So the net effect of the window in the spectrum analysis is to optimize the side lobes suppression and increase main lobe power level.

The wind profile detection of MST Radar signal is nothing but the measurement of Doppler of the signal due to scattering of the atmospheric elements. The returned signal from the atmospheric layers is very weak in terms of power. The received back-scattered signals called as Radar returns are associated with a lot of noise. The noise dominates the received signal as the distance between the Radar and the atmospheric layer targets increases and this leads to a decrease in Signal to Noise ratio(SNR). Hence the improvement of SNR of the atmospheric signals is investigated with the variable windows by varying the shape parameter of window.

The specifications of the MST radar data are given in Table1. The MST radar data obtained from the six directions are used to carry on the analysis. The complete implementation scheme to study the effect of shape parameter in different windows on the SNR of the radar returns is presented here.

- a) Compute the variable window with specified shape parameter.
- b) Taper the radar data with the window parameters specified.
- c) Perform the Fourier Transform on the tapered data.
- d) Compute the Signal to Noise Ratio (SNR)
- e) Compute the Mean Value of SNR Below zero (MVBZ)
- f) Compute the Mean Value of SNR Above Zero (MVAZ)
- g) Update value of shape parameter and repeat the steps (b)-(f).

Lower Stratosphere (up to 30 Km) -MST RADAR, Gadanki, India

No. Of Range Bins	150
No. of FFT points	512
No. of Coherent Integrations	64
No. of Incoherent Integrations	1
Inter Pulse Period	1000 μ sec
Pulse Width	16 μ sec
Beam	10
Period of Observation	July 2008
Pulse Width	16 μ s
Range resolution	150 m
Inter Pulse Period	1000 μ s
No of Beams	6 (E10y, W10y, Zy, Zx, N10x, S10x)
No of FFT points	512
No of incoherent integrations	1
Maximum Doppler Frequency	3.9 Hz
Maximum Doppler Velocity	10.94 m/s
Frequency resolution	0.061 Hz
Velocity resolution	0.176 m/s

E10y =East West polarization with off-zenith angle of 10^0
 W10y =East West polarization with off-zenith angle of 10^0
 N10x =North South polarization with off-zenith angle of 10^0
 S10x = North South polarization with off-zenith angle of 10^0

Table.1 Specifications of the MST Radar, India data on which the analysis is performed

4. DOLPH-CHEBYSHEV WINDOW ANALYSIS

The optimality criterion of the Dolph-Chebyshev window is that it's the Fourier transform exhibits main-lobe width narrowest for a selected/given side-lobe level. The Fourier transform of Dolph-Chebyshev window exhibits equal ripple for the specified side-lobe level. The Fourier Transform of this window is a mapping of algebraic Nth Chebyshev polynomial to the Nth order trigonometric Chebyshev polynomial by $T_N(X) = \cos(Nq)$. The Dolph-Chebyshev window is defined in terms of Fourier transform of uniformly spaced samples of the polynomial. These samples are defined in.

$$\left. \begin{aligned}
 W(k) &= (-1)^k \frac{\cosh \left[N \cosh^{-1} \left(\beta \cos \left(\pi \frac{k}{N} \right) \right) \right]}{\cosh [N \cosh^{-1}(\beta)]} : 0 \leq k < N-1 \\
 &\text{where } \beta \text{ is defined by :} \\
 \beta &= \cosh \left[\frac{1}{N} \cosh^{-1} \left(10^{\frac{-A}{20}} \right) \right] \\
 &\text{and where } A = \text{Sidelobe level (in dB)} \\
 w(n) &= \sum_{k=0}^{N-1} w(k) e^{j \frac{2\pi}{N} nk} : \text{ and where } W(N-k) = W(-k)
 \end{aligned} \right\} \dots\dots (4)$$

The Fourier Transform of Chebyshev window exhibits constant level side-lobes levels or uniform (inherited from the Chebyshev polynomial) and contains impulses in its time series. These impulses are located at the window boundaries in time domain. Figure 2 indicate the time and frequency description of a Dolph-Chebyshev window. The equal ripple behavior of the Dolph-Chebyshev window can also be obtained iteratively by the Remez (or the Equal-Ripple or Parks-McLellan) Filter design method.

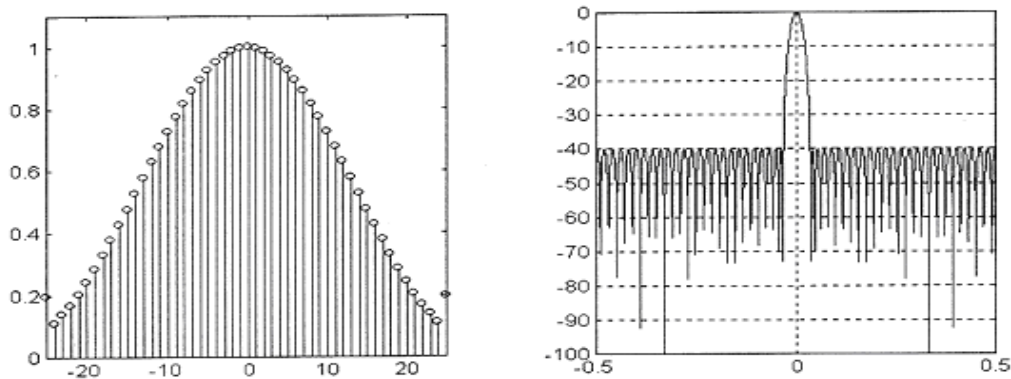


Fig.2 Dolph-Chebyshev (40dB) window and its Fourier Transform

SNR computation for the six sets of Radar data is carried using Dolph-Chebyshev window by varying the shape parameter “ α ” of the window and SNR variation with increase in “ α ” is presented in Fig.3. From Figs.3(a)-(f), in the case of East beam, SNRs (MVBZ) for the entire 150 bins taken into account, increases with the side lobe attenuation factor “ α ”. But in the case of North, South and West beams there is not much improvement is observed. In the case of Zenith-Y and Zenith-X beams, MVBZ increases with the side lobe attenuation of 20dB and decreases beyond 30dB. This can be attributed to the reason that the generation mechanism of the zenith beams is different from other beams. At the same time in all the Six-sets of data, the Mean value of the Above zero SNRs (MVAZ) increases with side lobe attenuation “ α ”. It attains a steady value when “ α ” is around 60dB.

From the Table (2) it is also observed that for the lowermost 50 bins, the MVBZ and MVAZ are not improved appreciably. In addition, a slight and marginal decrease in both SNR’s is observed. For the middle 50bins and the uppermost 50 bins the increase in MVBZ values is almost 3dB - 9dB when side lobe attenuation “ α ” reaches a value around 50dB-60dB. Further slight improvement is also seen when “ α ” is increased beyond 60dB. This result is important since the back-scattered signal from the middle and uppermost bins is very weak and improvement in SNR demands for the design of windows with good side lobe behavior for spectral estimation.

Noting the above observations, it is concluded that the Dolph-Chebyshev window can be used with “ α ” corresponding to the minimum of side lobe attenuation of 50dB to taper the data for spectral analysis. The results suggest that there is effect of side lobe reduction in improvement of SNR of data with noise and the design of optimal windows.

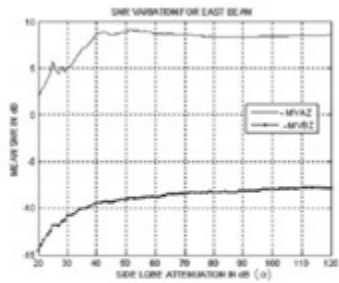


Fig.: 3.(a): AVERAGE SNR EAST Beam

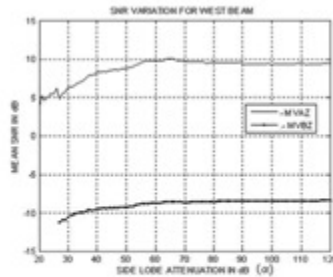


Fig.:3.(b): AVERAGE SNR WEST Beam

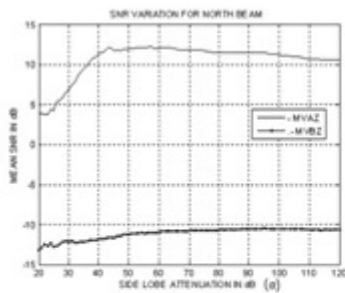


Fig.: 3.(c): AVERAGE SNR NORTH Beam

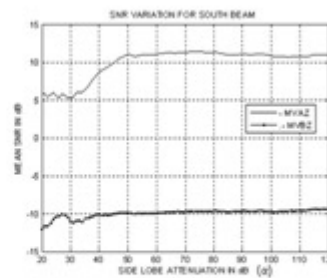


Fig.: 3(d): AVERAGE SNR SOUTH Beam

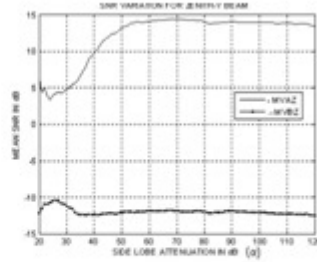
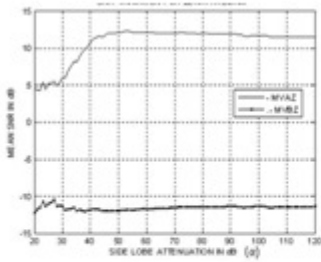


Fig.: 3.(e): AVERAGE SNR ZENITH-X Beam Fig.: 3.(f): AVERAGE SNR ZENITH-Y Beam

Beam	MVBZ			MVAZ		
	Lower bins	Middle bins	Upper bins	Lower bins	Middle bins	Upper bins
East	3	7	8	8	1	1
West	2	9	7	7	2	3
North	1	7	8	9	6	2
South	3	6	8	10	4	0
Zenith-x	4	10	5	10	2	0
Zenith-y	2	8	5	10	0	0

Table 2: Dolph Chebyshev Window SNR improvement in dB

5. KAISER WINDOW ANALYSIS

Another window designed in accordance with an optimality criterion, is the Kaiser-Bessel (or Prolate Spheroidal Wave Function). The previous window was characterized by minimum main-lobe width for a given side-lobe level and minimum bandwidth by approximately a minimum time-bandwidth product. Both windows had defects: one exhibited constant level side-lobes (resulting in high integrated side-lobes), the other exhibited excessive main-lobe width. An alternate, and related, optimality criterion is the problem of determining the wave shape on a finite support that maximizes the energy in a specified bandwidth. This wave-shape has been identified by Slepian, Landau, and Pollak as the Prolate Spheroid function (of order zero) which contains a selectable time-bandwidth product parameter. Kaiser has developed a simple numerical approximation to this function in terms of the zero order modified Bessel function of the first kind (so the designation Kaiser- Bessel). The Kaiser-Bessel window is defined in (6) where the

parameter α is the window's half time-bandwidth product. The series for the Bessel function converges quite rapidly due to the $k!$ in the denominator.

The transform of the Kaiser-Bessel window is (within very low level aliasing terms) the function shown in (7), We see that this function tends to $\frac{\sin(x)}{(x)}$ when the spectral argument is evaluated beyond the time-bandwidth related main-lobe bandwidth.

It must be noted that the Kaiser window can also be approximated by samples of the main-lobe of its spectra since the window is self-replicating under the tune-limiting and band-limiting operations.

Figure 6 presents the Kaiser-Bessel window for parameter selected ' γ ' to achieve 60 dB side-lobes.

$$w(n) = \frac{I_0 \left[\pi\alpha \sqrt{1.0 - \left(\frac{n}{N/2} \right)^2} \right]}{I_0[\pi\alpha]} \quad \text{where} \quad \dots\dots\dots (5)$$

$$I_0(x) = \sum_{k=0}^{\infty} \left[\frac{(x/2)^k}{k!} \right]^2$$

$$W(\theta) = \frac{N}{I_0(\alpha\pi)} \frac{\sinh \left[\sqrt{(\alpha\pi)^2 - (N\theta/2)^2} \right]}{(\alpha\pi)^2 - (N\theta/2)^2} \quad \dots\dots\dots (6)$$

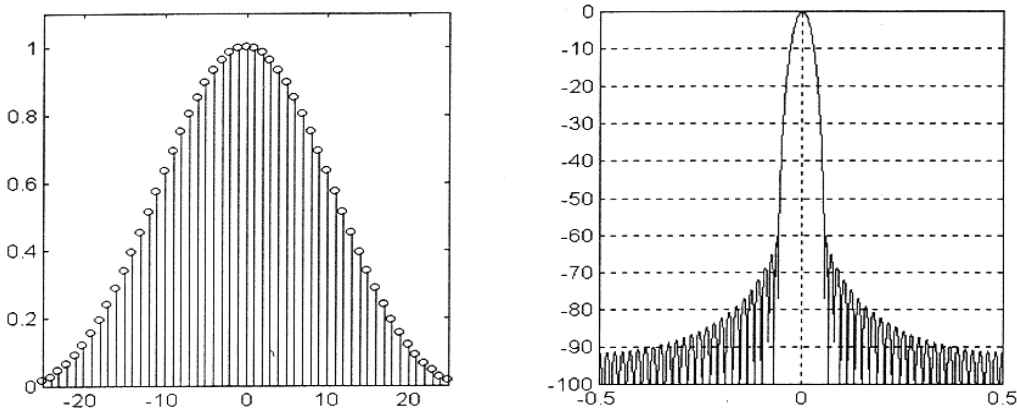


Figure: 6 Kaiser-Bessel (60dB, $\alpha=8.3$) window and its Fourier Transform.

The SNR is computed for all the six sets of Radar data and is presented in Figure 7(a)-(f). It can be inferred that when $\beta=0$, which corresponds to rectangular window, the average lower SNR than zero (MVBZ) is lowest and increases with increase in parameter β steadily. When β is above 4 (MVBZ) increases slowly and reaches to a steady value. Further increase of β produces not much appreciable increase in average SNR value of MVBZ. On the other in all the cases of six sets of data, the average positive SNR (MVAZ) has not suffered any major difference when β is increased. Moreover, a slight decrease in average positive (MVAZ) SNR is observed when β is beyond 6.

From the Table 3 it is seen that for the lowest 50 bins, the average SNRs in the negative and positive sides (MVBZ and MVAZ) are not improved appreciably. In addition there is slight decrease in both SNRs are observed. For the middle 50 bins and the upper 50 bins the increase in average negative SNR (MVBZ) values is almost 5 dB to 6 dB when β is around 6. Further improvement is seen when the β is increased beyond 6. This result is important since the reflected signal from the middle and uppermost bins is very weak and improvement in SNR is highly desirable in estimating the spectrum. With these observations, it is concluded that the Kaiser window can be used with β greater than or equal to 6 to taper the data for spectral analysis in place of a normal rectangular window. This result also suggests that, the effect of side lobe reduction enhances in the improvement of SNR of noisy data and suggests for the design of optimal windows.

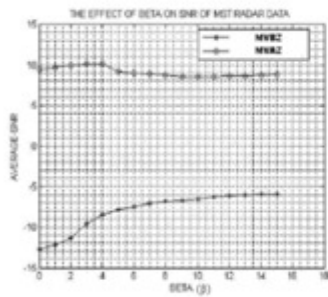


Fig.: 7(a): AVERAGE SNR EAST Beam

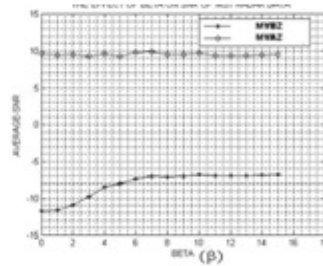


Fig.: 7(b): AVERAGE SNR WEST Beam

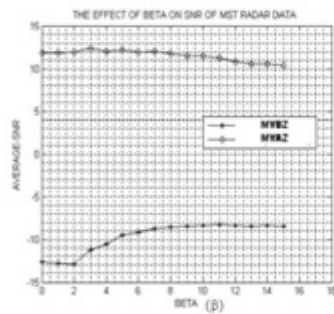


Fig.: 7(c): AVERAGE SNR NORTH Beam

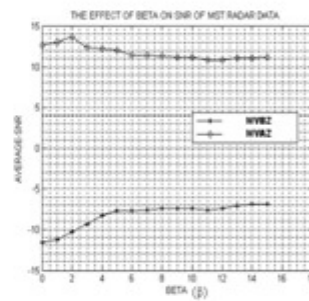


Fig.: 7(d): AVERAGE SNR SOUTH Beam

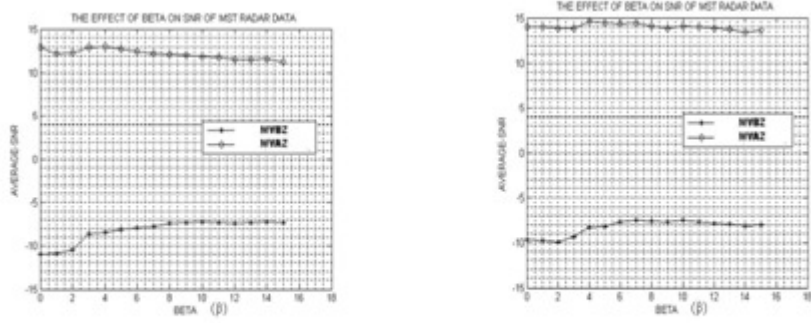


Fig.: 7(e): AVERAGE SNR ZENITH-X Beam Fig.:7(f): AVERAGE SNR ZENITH-Y Beam

Table 3: Kaiser Window SNR improvement in dB

Beam	MVBZ			MVAZ		
	Lower bins	Middle bins	Upper bins	Lower bins	Middle bins	Upper bins
East	3	4	7	0.5	0.2	2
West	0	3	8	0.2	0.1	2
North	1	4	6	0	0.1	1
South	0	3	5	0.1	0.1	0
Zenith-x	0	4	4	0	0.1	0.1
Zenith-y	0	2	3	2	0	0.1

6. GAUSSIAN WINDOW

A second window that exhibits a measure of optimality is the Gaussian or Weierstrass function. A desired property of a window is that there are smooth (usually) positive functions with Fourier Transforms which approximate an impulse, (i.e tall thin main lobe with low level side-lobes).

The sampled Gaussian window is defined in (7) with the parameter ' γ ', the inverse of the standard deviation, controlling the effective time duration and the effective spectral width.

$$w(n) = \exp \left[-\frac{1}{2} \left(\gamma \frac{n}{N/2} \right)^2 \right] \quad \dots\dots\dots (7)$$

The Fourier Transform of this truncated window is the convolution of the Gaussian transform with a Dirichlet kernel . The convolution results in the formation of the spectral main-lobe (approximating the target's main-lobe) with accompanying side-lobes whose peak levels depend on the parameter ‘ γ ’. As expected, larger ‘ γ ’ leads to wider main-lobe and lower side-lobes. Figures 3 present Gaussian windows with parameter ‘ γ ’ selected to achieve 60 dB side-lobe level. It is noted that the main-lobes are considerably wider than those of the Dolph-Chebyshev and the upcoming Kaiser-Bessel windows. A useful observation made is that the main-lobe of the Gaussian window is 1/3 again wider than the Blackman-Harris window exhibiting the same side-lobe level.

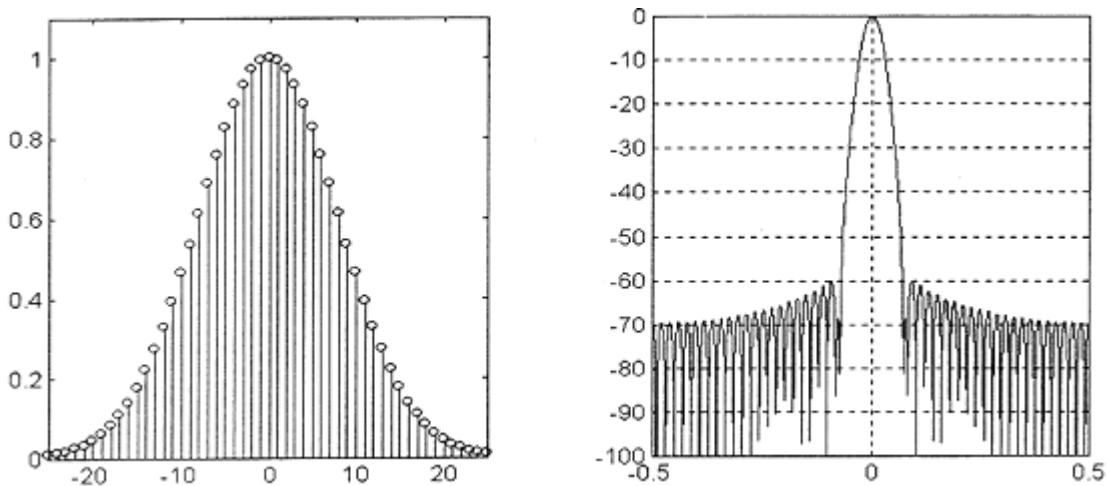


Fig.8 Gaussian (60 dB, $\alpha=3.1$) window and its Fourier Transform.

The SNR computation is done for the six sets of Radar data and is presented in Figure.9 (a)-(f). The side lobe attenuation factor ‘ γ ’ is varied from 1 to 10. From Figure 9 (a)-(f) it is observed that in the case of East beam, West beam, North beam, South beam, Zenith-X the Mean value of the below zero SNRs (MVBZ) for the entire 150 bins taken into account, increases with the side lobe attenuation factor ‘ γ ’. But in the case of Zenith-Y beam, there is no appreciable change observed. In the case of East, West, North and South beams MVBZ increases when the side lobe attenuation is in between 1dB and-4dB and remains constant beyond 4dB. This is attributed to the truth that the generation mechanism of the beams of zenith is different. On the other hand in all the Six-sets of data, the Mean of the Above zero SNRs (MVAZ) marginally reduces with side lobe attenuation γ . It attains a steady value when ‘ γ ’ is in between 5dB and 10dB.

From the Table.4 it is also observed that for the down most 50 bins, the MVBZ and MVAZ are not improved appreciably. In addition, a slight/ marginal decrease in both SNRs is observed. For the middle 50 bins and the uppermost 50 bins, there is increase in MVBZ and is almost 2dB- 8dB when side lobe attenuation γ varies from 1 to 10. Further, slight improvement is also seen when ‘ γ ’ is increased beyond 2 and above. This result is important since the reflected/back-scattered

signal from the middle and uppermost bins is weak and improvement in SNR is highly required in spectral estimation. Noting the above observations, it may be concluded that the Gaussian window can be used with “ γ ” corresponding to the minimum of side lobe attenuation of 2 to taper the data for spectral analysis. The results indicates the effect of side lobe level reduction in the improvement of the SNR of noisy signal.

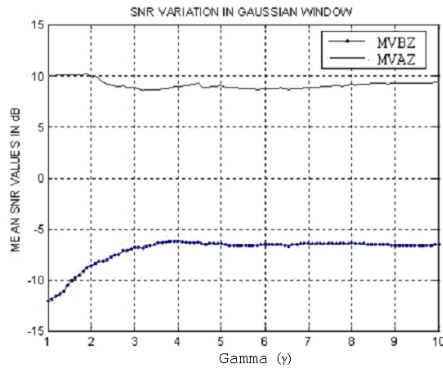


Figure.9.1(a): AVERAGE SNR EAST Beam

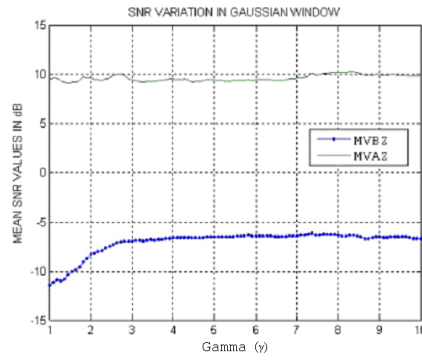


Figure 9.1(b): AVERAGE SNR WEST Beam

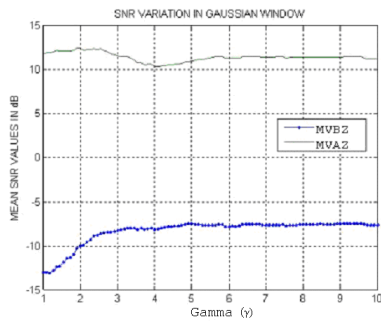


Figure.9.1(c): AVERAGE SNR NORTH Beam

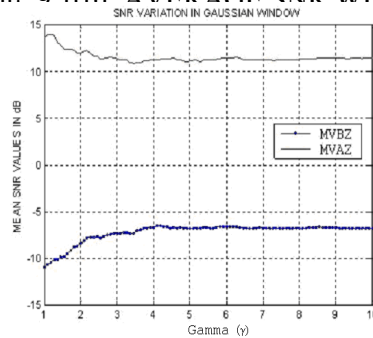


Figure.9.1(d): AVERAGE SNR SOUTH Beam

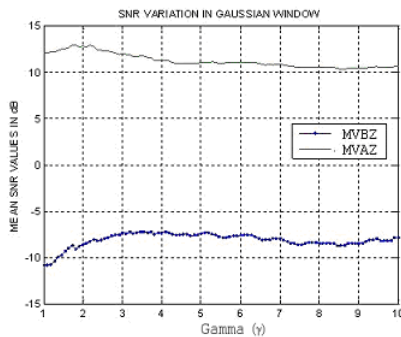


Figure.9.1(e): AVERAGE SNR ZENITH-X Beam

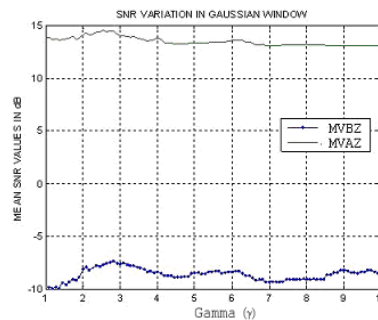


Figure.9.1(f): AVERAGE SNR ZENITH-Y Beam

Beam	MVBZ			MVAZ		
	Lower bins	Middle bins	Upper bins	Lower bins	Middle bins	Upper bins
East	2	6	8	1	3	6
West	2	5	7	2	4	4
North	6	7	8	1	5	6
South	6	5	6	1	5	2
Zenith-x	5	6	7	1	5	2
Zenith-y	2	4	4	1	1	3

Table 4: Gaussian Window SNR improvement in dB

7. TUKEY WINDOW

Numerous investigators have constructed windows as products, as sums, as sections, or as convolutions of simple functions of other simple windows. These windows have been constructed for certain desirable features. The Tukey window, often called the *cosine tapered window*, is best imagined as a cosine lobe of width $(\alpha/2) N$ convolved with a rectangle window of width $(1.0-\alpha/2) N$. Of course the resultant transform is the product of the two corresponding transform. The window represents an attempt to smoothly set the data to zero at the boundaries while not significantly reducing the processing gain of the windowed transform. The window evolves from the rectangle to the Hanning window as the parameter α varies from zero to unity. The family of windows exhibits a confusing array of sidelobe levels arising from the product of the two components transforms.

The window is defined by

$$w(n) = \begin{cases} 1.0, & 0 < |n| < \alpha \frac{N}{2} \\ 0.5 \left[1.0 + \cos \left[\pi \frac{n - \alpha \frac{N}{2}}{2(1 - \alpha) \frac{N}{2}} \right] \right], & \alpha \frac{N}{2} < |n| < \frac{N}{2} \end{cases} \dots\dots\dots (8)$$

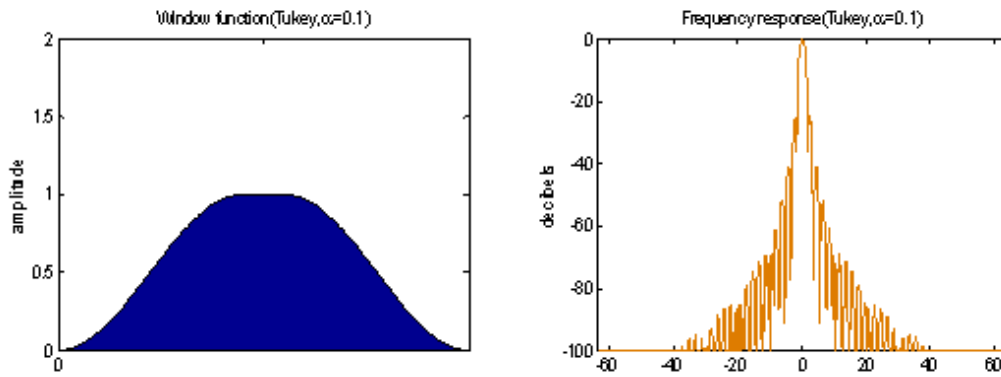


Figure10: Tukey window and Fourier transform

The SNR computation for the six sets of Radar data is carried on and presented in Figure.11(a)-(f) and Table 5. From Figure 11(a)-(f) it is observed that in the case of East beam, West beam, North beam, South beam, Zenith-X the Mean value of the below zero SNRs (MVBZ) for the entire 150 bins taken into account, increases with the shape factor α . But in the case of Zenith-Y beam, there is no appreciable change observed “ α ”. In the case of East, West, North and South beams MVBZ increases when the shape factor is in between zero and 0.8 and remains constant beyond 1. This may be attributed to the fact that the generation mechanism of the zenith beam is different. On the other side in all the cases of six-sets of data, the Mean value of the Above zero SNRs (MVAZ) reduces marginally with “ α ”.

From the Table 5 it is also observed that for the lowermost 50 bins, the MVBZ and MVAZ are not improved appreciably. Moreover, a slight/marginal decrease in both types SNRs is observed. For the middle 50 bins and the uppermost 50 bins, the increase in MVBZ is almost 2dB- 8dB when side lobe attenuation varies from 1 to 10. This result is important since the back-scattered signal from the middle and uppermost bins is very weak and improvement in SNR is highly desirable in spectral estimation. It is observed that the Tukey window can be used with “ α ” corresponding to the minimum of side lobe attenuation of 18dB to taper the data for spectral analysis. From the results, it is noted that there is effect of side lobe reduction in the improvement of SNR of noisy data/signal. For ‘ α ’ greater than 0.8 the SNR improvement is good at higher range bins. Based on the above results shows that maximum side lobe level from the Table 5, it is observed that SNR improvement is achieved with the windows having variable parameters producing the side lobe level is less than 20 dB.

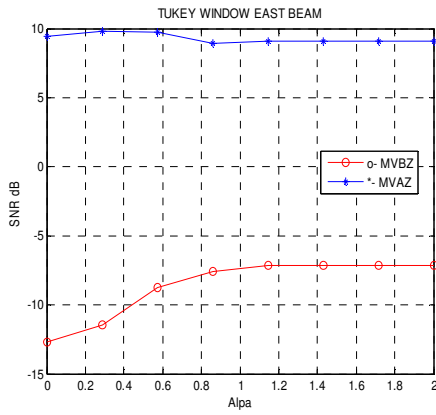


Fig. 11 (a): Tukey window East Beam

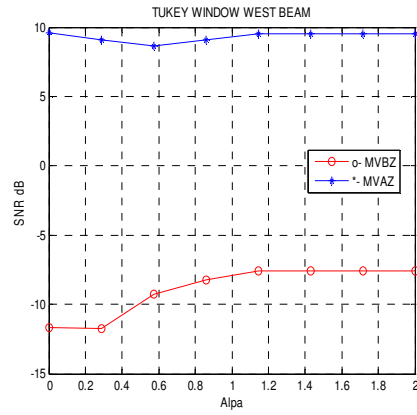


Fig.11 (b): Tukey window West Beam

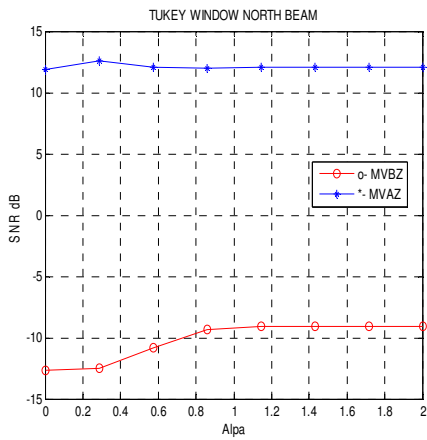


Fig. 11 (c): Tukey window North Beam

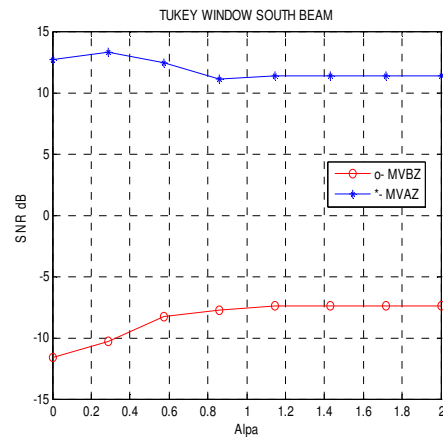


Fig. 11 (d): Tukey window South Beam

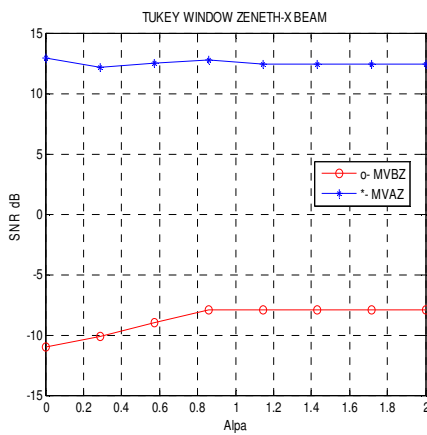


Fig.11(e): Tukey window Zenith-X Beam

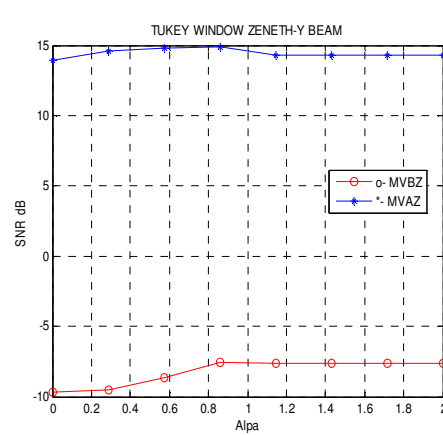


Fig.11(f): Tukey window Zenith-Y Beam

Beam	MVBZ			MVAZ		
	Lower bins	Middle bins	Upper bins	Lower bins	Middle bins	Upper bins
East	2	3	8	1	0	2
West	0	3	7	0	4	5
North	1	4	5	0	4	4
South	0	4	7	0	0	0
Zenith-x	0	3	3	0	0	0
Zenith-y	0	3	2	1	0	0

Table 5: Tukey Window SNR improvement in dB

8. DISCUSSION AND CONCLUSIONS

It is a well known fact that for the lower bins, the SNR will be positive in most of the cases, where as the back scattered signal from upper bins is very weak, the SNR values are negative. Similarly for the middle bins, the SNR will be in between these two cases. In this study the 150 bins of the MST radar data is divided into three equal parts viz. lower bins (1-50), middle bins (51-100) and upper bins (101-150) and the SNR variations are studied separately in all these three regions. The positive average SNR values are termed as Mean Value Above Zero (MVAZ) and the negative average SNR values are termed as Mean Value Below Zero (MVBZ) and the same is followed throughout this work in interpreting the results. The noise dominates the signal as the distance between the Radar and target increased and this leads to a decrease in Signal to Noise ratio. Hence it is highly necessary to improve the SNRs of the upper bins, which is clearly observed in the results.

In all these cases observed above, it is seen that after certain value of the optimum constants viz. α , β , γ of the adjustable parameter windows, there is no appreciable change in the improvement of SNR for the upper bins. This can be attributed by the fact that the side lobe reduction in windows improves the SNR values. But at the same time the side lobe reduction also increases the main lobe width which in turn increases the spectral leakage which directly contributes to the decrease in the SNR values.

Therefore, the improvement in SNR due to the increase in side lobe attenuation will be compensated by the increase of main lobe width with no further improvement in SNR beyond certain value of the optimum constants is observed in the present study. This is particularly important to note that theoretical fact that good windows are the optimization of the side lobe attenuation and main lobe width and the same is proved for the practical MST radar data analysis.

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